

# How to Relate Particle Physics and Air Shower Development : the EPOS Model

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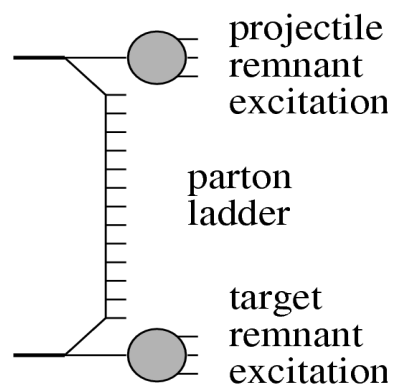
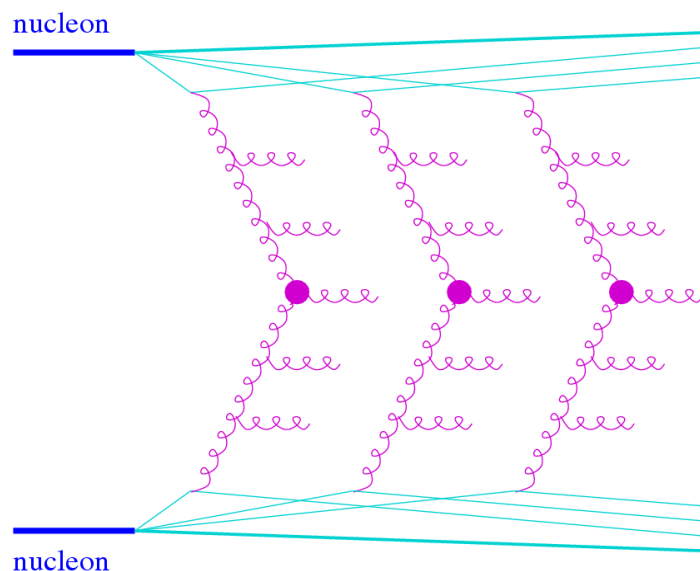
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# Outline

- **The EPOS model**
- **Constrains from air showers**
  - ➔ Cross section
  - ➔ Multiplicity
  - ➔ Forward spectra
  - ➔ Baryon production
- **New results**

# The EPOS Model



**EPOS\* is a parton model, with many binary parton-parton interactions, each one creating a parton ladder.**

- ➔ Energy-sharing : for cross section calculation AND particle production
- ➔ Parton Multiple scattering
- ➔ Outshell remnants
- ➔ Screening and shadowing via unitarization and splitting

**EPOS designed to be used for particle physics experiment analysis (SPS, RHIC, LHC)**

**EPOS is a complex monte-carlo for minimum bias hadronic interaction generation (h-p to A-B) : suitable for air shower !**

# EPOS and Air Showers

## Success :

- ➔ Large increase of # muons (MIA data, Auger S1000)
  - baryon-antibaryon contribution
  - remnant breakup

## Problems :

- ➔ KASCADE data (talks by Haungs and Hörandel) :
  - composition too light : too many muons  
or not enough elec. ?
  - Hadron energy too low : shower too old ?
- ➔ Auger : not enough muons in inclined showers :
  - Muon energy too low : shower too old ?

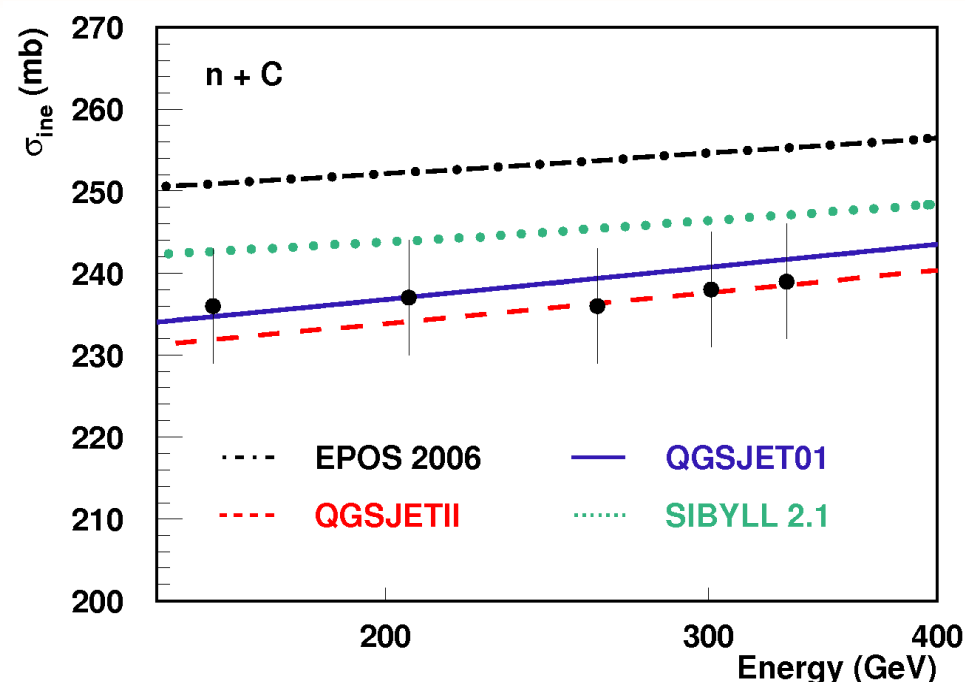
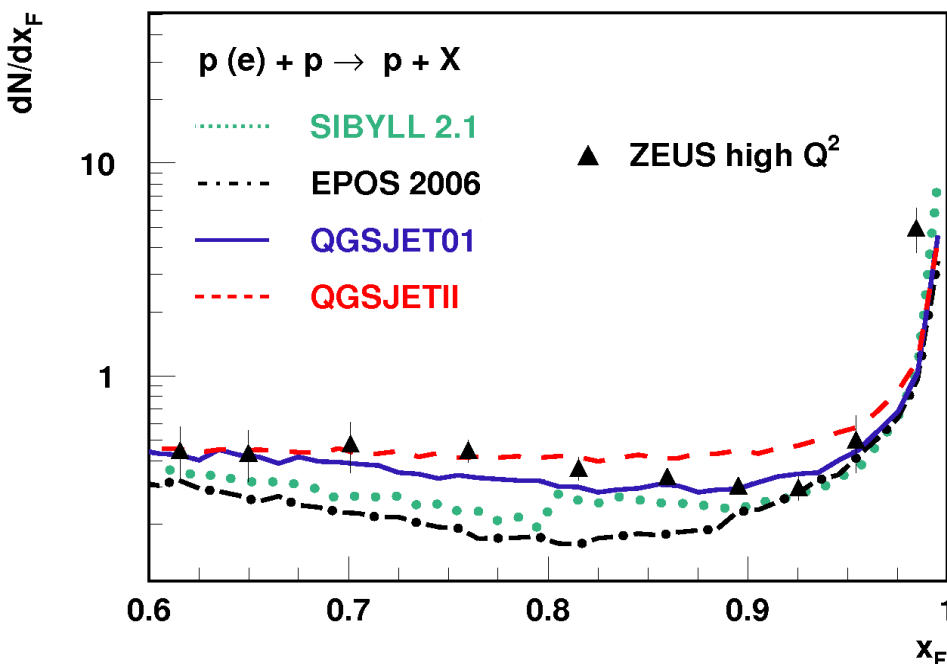
needed here !

cross section  
or inelasticity  
too large ?

# Cross Section and Inelasticity

## Cross-section too large in pA :

➔ Problem with the calculation of pA cross- sections



Number of protons at large  $x_F$  too small :

➔ remnant break-up too large at high energy  
➔ inelasticity too large

**Problem seen in CR confirmed by accelerator data !**

# Constrains by Air Showers

## Air shower data give global constrains on the hadronic interactions

### ● Low cross section / inelasticity

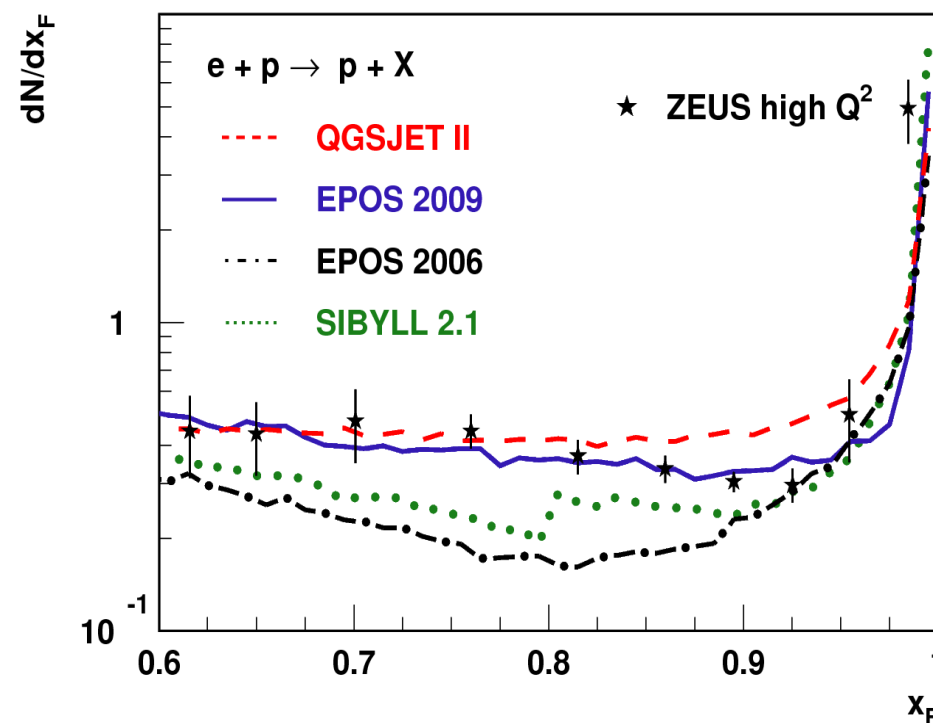
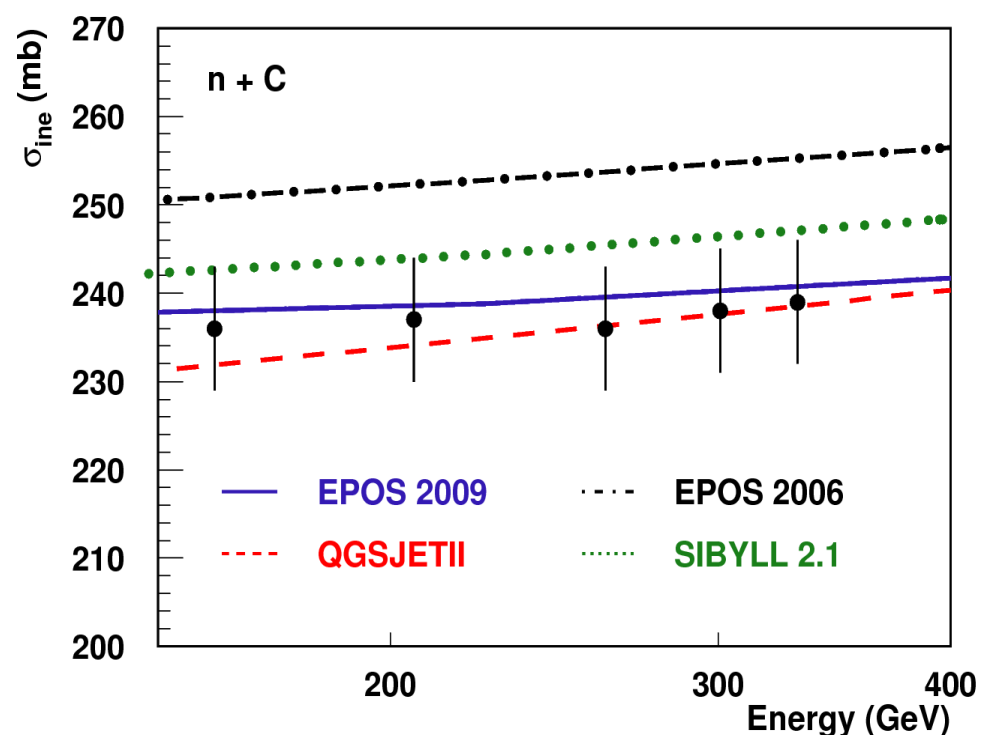
- ➔ Equilibrium remnant/central particle production
- ➔ Strong screening effect but with saturation
  - Reduction of multiple interaction
  - Limited growth of multiplicity
  - Limited growth of hard scattering ???

### ● High muon number

- ➔ Modified fragmentation : evolution with energy
  - Increase of diquark production
  - Cronin effect (pt kink)
  - $\pi^0$  spectra

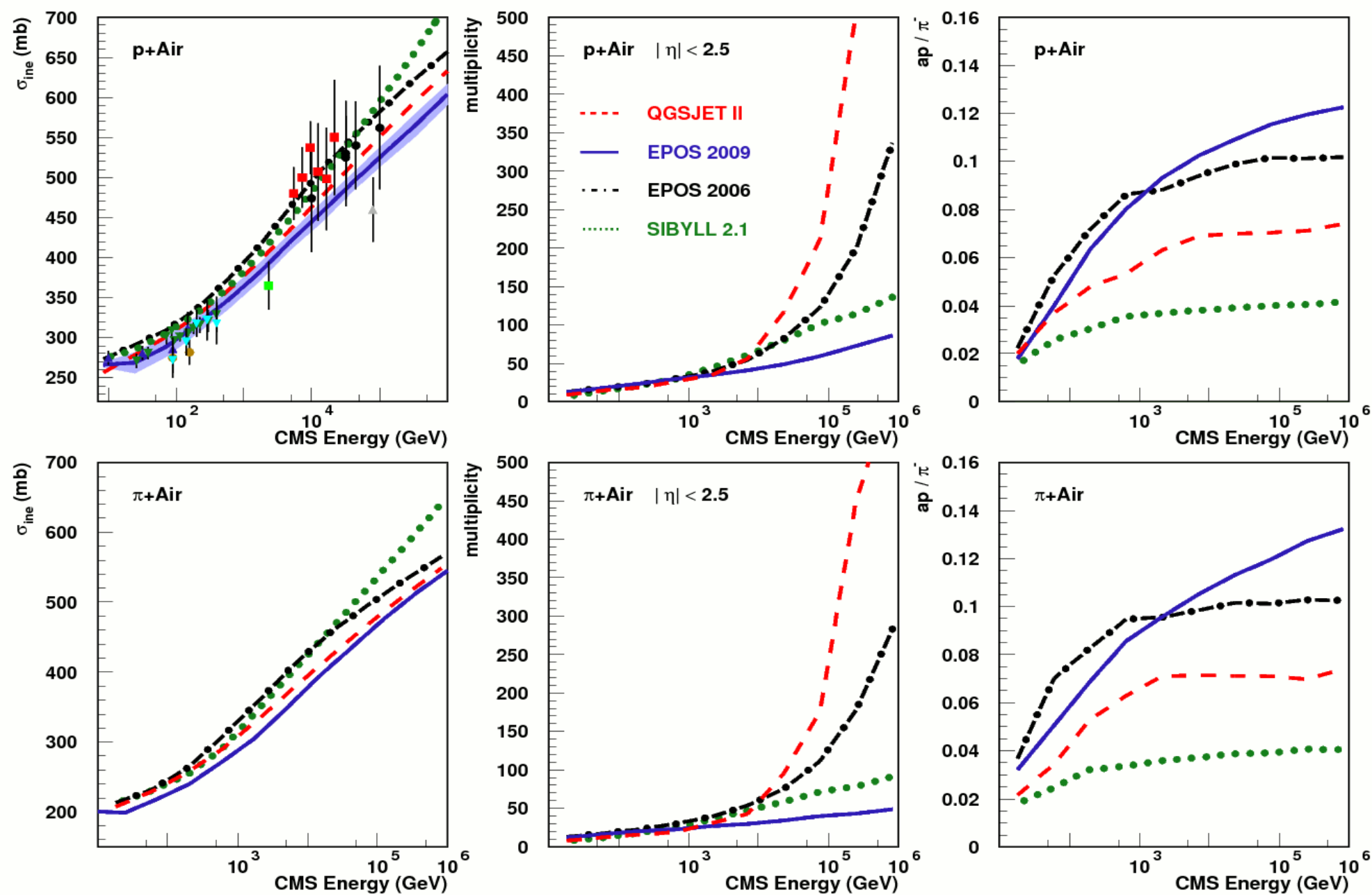
# New Developments

- To increase predictive power : less parameters
  - ➔ improvement of screening effect
  - ➔ unique saturation scale (pA cross section and RHIC  $p_t$ )
- More constraints on other parameters (break-up, baryons, ...)



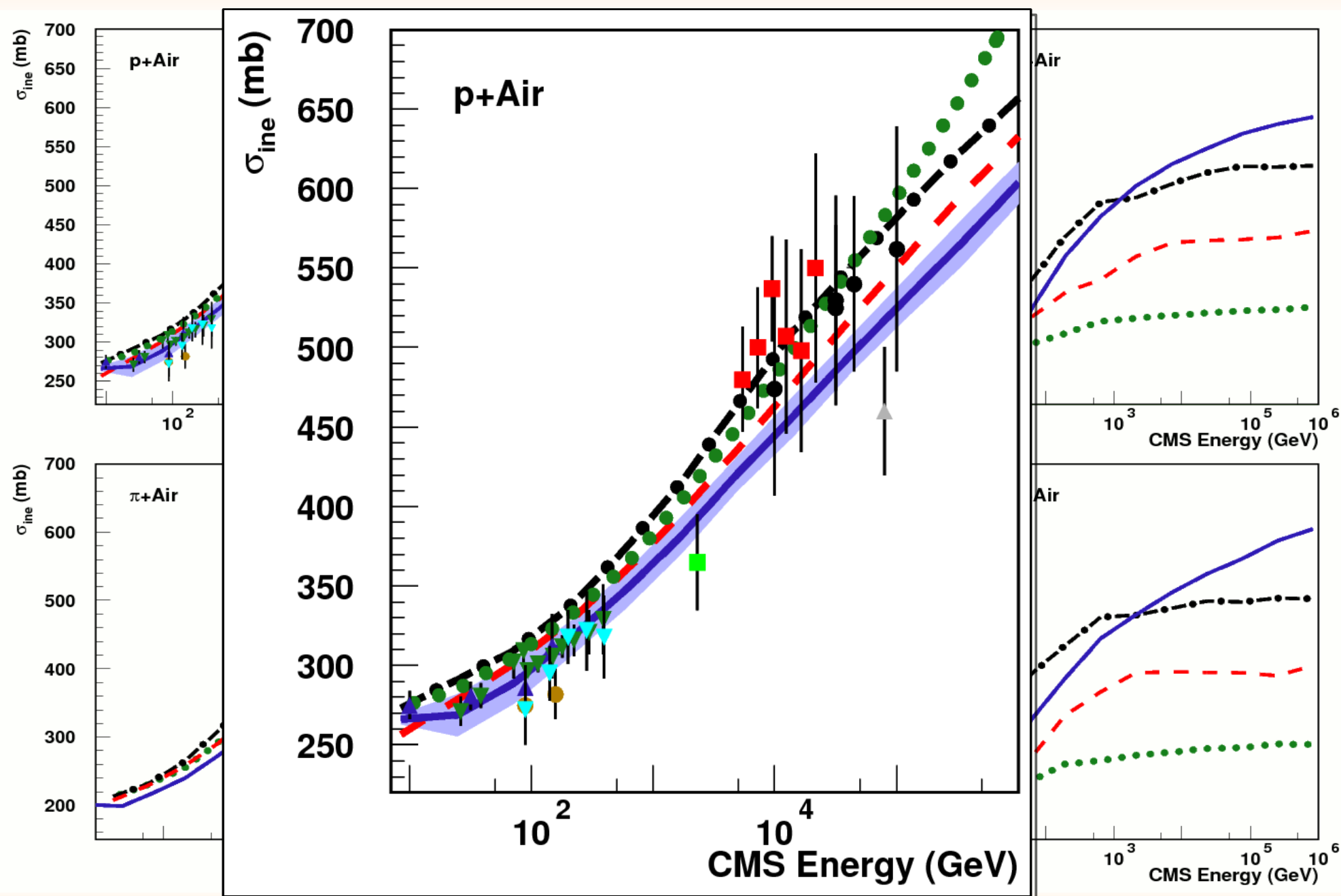
EPOS 2009 = EPOS 1.99 last release, EPOS 2006 = EPOS 1.61 prev. release

## h-Air Results



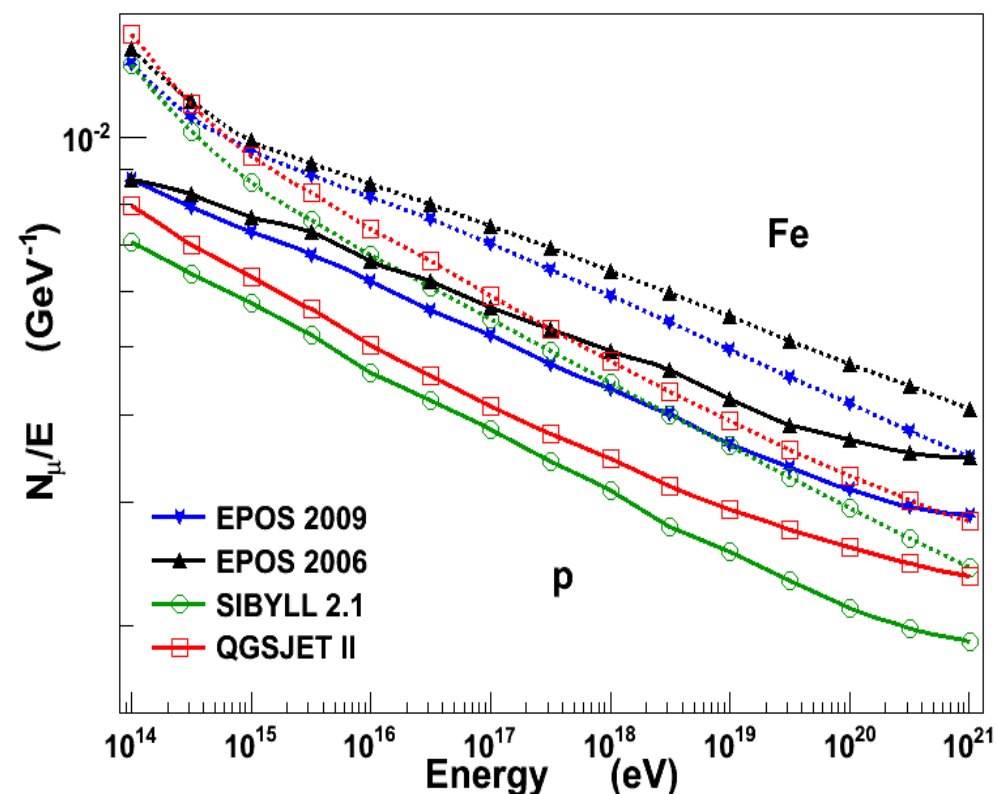
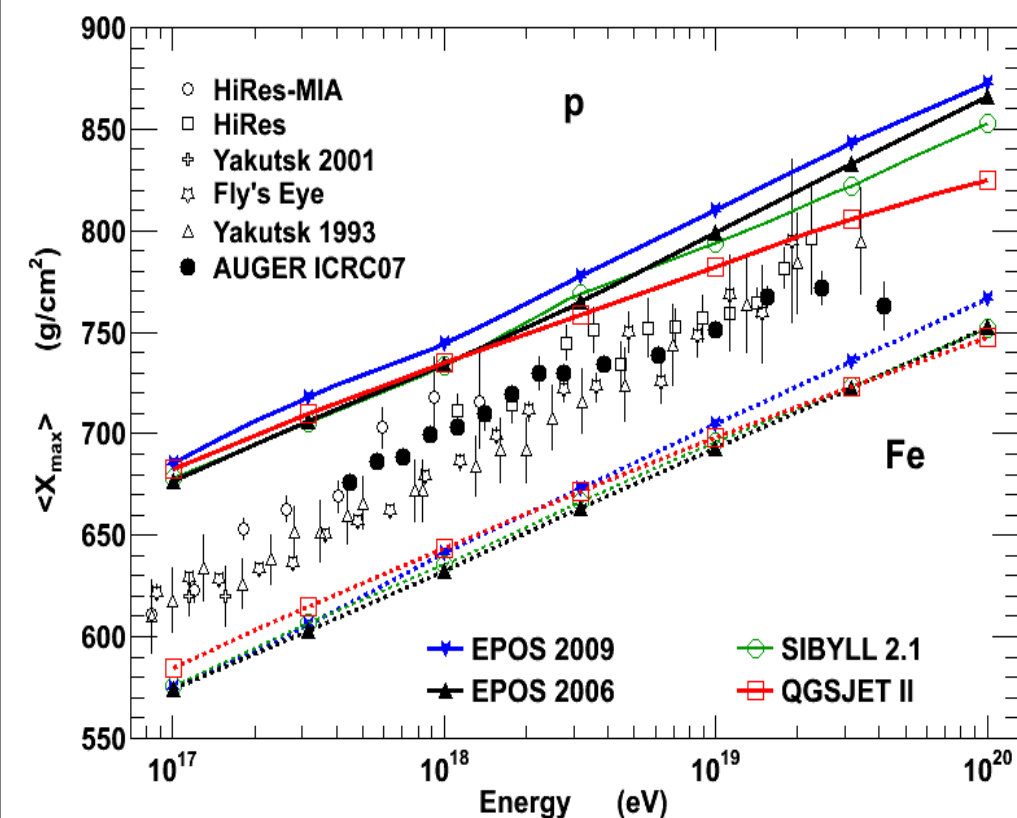


# h-Air Results

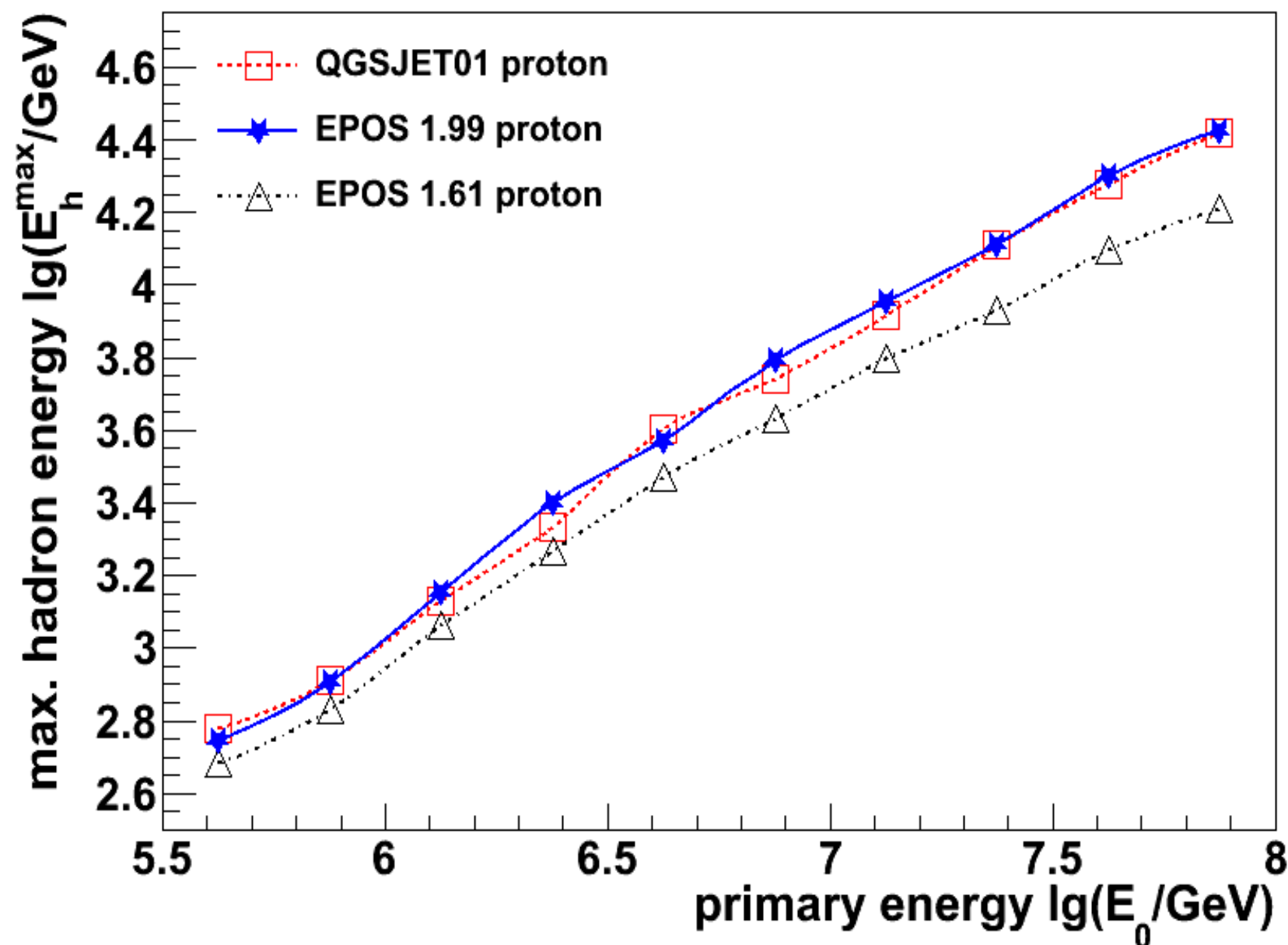


# Consequences on Air Showers

- Smaller cross section = deeper shower (larger  $X_{\max}$ )
  - More electron @ ground
  - More energetic hadrons @ ground
- Slightly less muons (less break-up) but with more energy



# Consequences on Air Showers



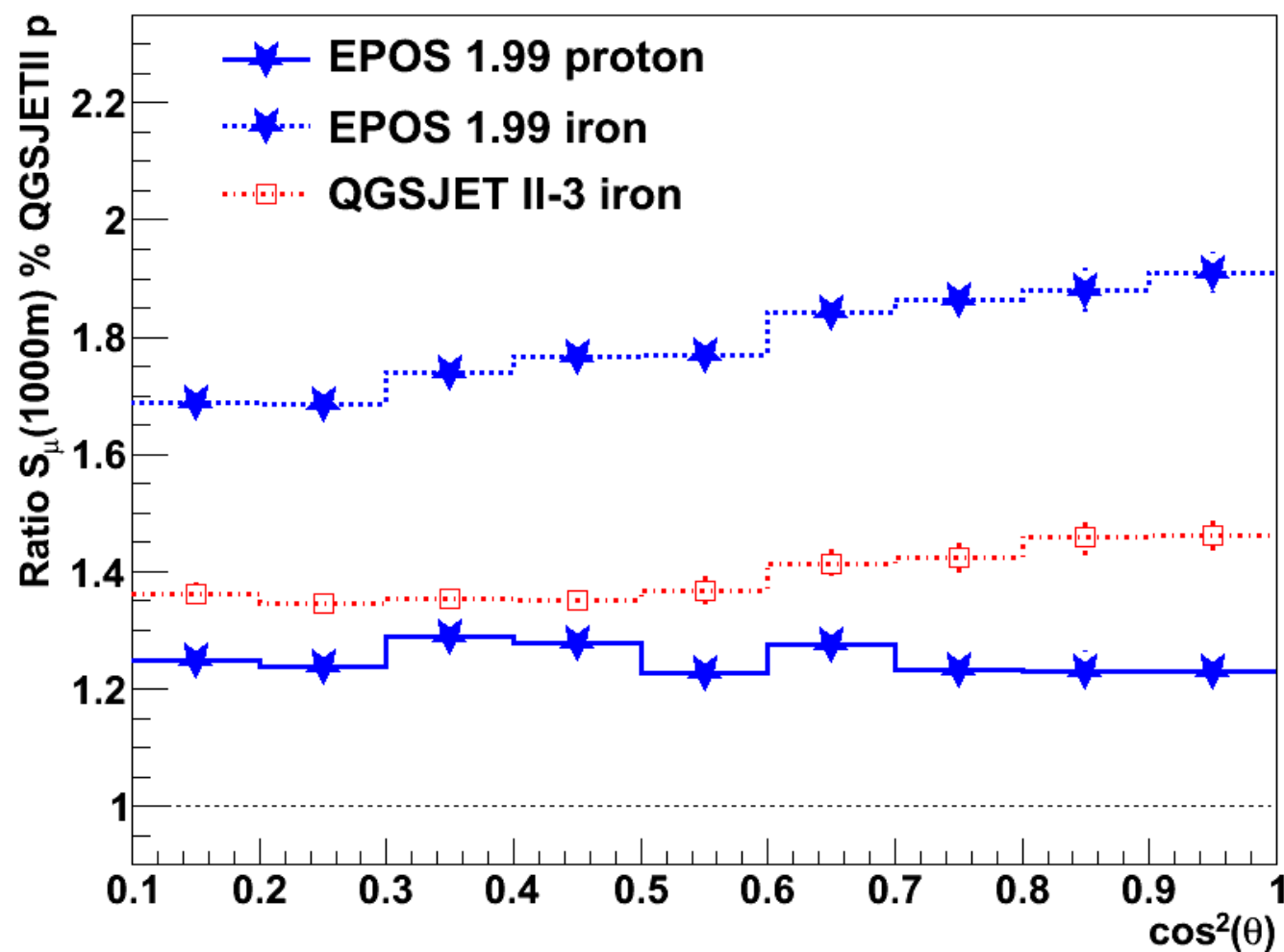
# Summary

## EPOS ... or how to relate cosmic ray physics and particle physics :

- ➔ **Particle physics** : baryons should be treated correctly
  - effects on muon number
- ➔ **Cosmic ray** : if a lot of muons, then pA cross section small
  - large nuclear screening effects
  - effect on multiplicity and  $p_t$  in pA
    - limitation by RHIC data on  $p_t$
    - cross section limited by nuclear saturation
- ➔ Low multiplicity and low cross section = large  $X_{\max}$  for Auger

**Consequences on cosmic ray mass composition**

# Muons @ 1000 m – $10^{19}$ eV



PAO observable